# Refined position estimates for surfaces of Willmore type in Riemannian manifolds 

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#### Abstract

In this paper we consider surfaces which are critical points of the Willmore functional subject to constrained area. In the case of small area we calculate the corrections to the intrinsic geometry induced by the ambient curvature. These estimates together with the choice of an adapted geometric center of mass lead to refined position estimates in relation to the scalar curvature of the ambient manifold.


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## 1. Introduction

Let $(M, g)$ be a three dimensional Riemannian manifold and let $\Sigma \subset M$ be a smooth, compact, two-sided, immersed surface. The Willmore energy of $\Sigma$ is defined as

$$
\mathcal{W}(\Sigma)=\frac{1}{4} \int_{\Sigma} H^{2} \mathrm{~d} \mu
$$

where $H$ is the mean curvature of the immersion and $\mathrm{d} \mu$ denotes the induced surface measure on $\Sigma$. We consider surfaces $\Sigma$ which are critical points of $\mathcal{W}$ subject to the constraint of prescribed area $|\Sigma|$. These surfaces satisfy the Euler-Lagrange equation

$$
\begin{equation*}
\Delta H+H|\AA|^{2}+H \operatorname{Ric}(\nu, \nu)+\lambda H=0 \tag{1}
\end{equation*}
$$

where $\lambda \in \mathbf{R}$ is the Lagrange parameter, $\Delta$ is the Laplace-Beltrami operator of the induced metric $\gamma$ on $\Sigma, \AA=A-\frac{1}{2} H \gamma$ is the trace free part of the second fundamental form $A$, and $\nu$ denotes (one choice of) the normal vector to $\Sigma$. Furthermore, Ric is the Ricci curvature of ambient metric $g$.

Concerning the existence of minimizers of the area constrained problem in compact manifolds, we have

Theorem 1.1. Let $(M, g)$ be a three dimensional Riemannian manifold. Then there exists $a_{\text {min }} \in(0, \infty)$ and for each $a \in\left(0, a_{\text {min }}\right)$ there exists $a$ smooth closed embedded surface $\Sigma_{a}^{m i n}$ such that

$$
\mathcal{W}\left(\Sigma_{a}^{\min }\right)=\inf \{\mathcal{W}(\Sigma) \mid \Sigma \text { smooth closed immersion and }|\Sigma|=a\}
$$

and $\left|\Sigma_{a}^{m i n}\right|=a$.
This was shown by Chen and $\mathrm{Li}\left[1\right.$ in the class of $W^{2,2}$-conformal immersions and by Lamm and the author [7] as well as by Mondino and Rivière [11] with the additional assertion of smoothness of the minimizing surfaces.

Critical points for this minimization problem can be constructed by perturbing geodesic spheres centered at a non-degenerate point of the scalar curvature. Independently Ikoma, Malchiodi, and Mondino [4] as well as Lamm, Schulze and the author [9] have shown the following:

Theorem 1.2. Let $(M, g)$ be a three dimensional Riemannian manifold and let $p \in M$ be such that $\nabla \mathrm{Sc}(p)=0$ and such that $\nabla^{2} \mathrm{Sc}(p)$ is non-degenerate. Then there exist $a_{\text {pert }} \in(0, \infty)$, a neighborhood $U$ of $p$, and for each $a \in$
( $0, a_{\text {pert }}$ ) a spherical surface $\Sigma_{a}^{\text {pert }}$ which satisfies (1) for some $\lambda \in \mathbf{R}$ and $\left|\Sigma_{a}\right|=a$. The $\Sigma_{a}$ are mutually disjoint and $\bigcup_{\left(0, a_{\text {pert }}\right)} \Sigma_{a}=U \backslash\{p\}$.

The shape and position of critical points, that is solutions to equation (1) was studied by Lamm and the author [6, 7] and with more general assumptions by Laurain and Mondino [10]. The position estimates implied by combining these three papers are the following:

Theorem 1.3. Let $(M, g)$ be a three dimensional Riemannian manifold. Then there exist $a_{0} \in(0, \infty)$ and constants $C_{1}, C_{2} \in(0, \infty)$ with the following property. Let $\Sigma \subset M$ be a surface satisfying equation (1) for some $\lambda \in \mathbf{R}$ such that $|\Sigma| \leq a_{0}$ and $\mathcal{W}(\Sigma)<4 \pi+a_{0}$. Then $\operatorname{diam}(\Sigma) \leq C_{1}|\Sigma|^{1 / 2}$ and $|\nabla \mathrm{Sc}| \leq C_{2}|\Sigma|^{1 / 2}$ on $\Sigma$.

A consequence of this Theorem is that the surfaces $\Sigma_{a}^{\min }$ concentrate near critical points of the scalar curvature of $M$. From the expansion of the Willmore functional in [6, Theorem 5.1] it follows in addition that the minimizers $\Sigma_{a}^{\text {min }}$ concentrate near the points in $M$ where the scalar curvature is maximal as $a \rightarrow 0$.

The previously cited results were to a large extend based on the observation, that surfaces satisfying (1) with small area and Willmore energy close to $4 \pi$ behave like their Euclidean counterparts. The aim of this paper is to provide a more precise description of the shape and position of solutions to (11), that take into account the perturbations induced by the ambient geometry.

As an application of the estimates we derive an improved position estimate.

Theorem 1.4. Let $(M, g)$ be a three dimensional Riemannian manifold with $C_{B}$-bounded geometry. Then there exists $a_{0} \in(0, \infty)$ and a constant $C \in(0, \infty)$ with the following property. For every surface $\Sigma \subset M$ satisfying equation (1) for some $\lambda \in \mathbf{R}$ with $|\Sigma| \leq a_{0}$ and $\mathcal{W}(\Sigma)<4 \pi+a_{0}$ there exists a point $p_{0}$ contained in the region enclosed by $\Sigma$ such that for all $p \in \Sigma$ we have $\operatorname{dist}\left(p_{0}, p\right)<\frac{3}{4} \operatorname{diam}(\Sigma)$ and $\left|\nabla \operatorname{Sc}\left(p_{0}\right)\right| \leq C|\Sigma|$.

For the definition of $C_{B}$-bounded geometry, refer to definition 2.1 .
The main use for this improvement of Theorem 1.3 is to further narrow down the position of the surfaces $\Sigma$ as in the theorem. To this end assume that the critical points of the scalar curvature of $(M, g)$ are such that the Hessian there is non-degenerate, in other words, that the scalar curvature
on $M$ is a Morse function. Let $\bar{p}$ be a critical point for Sc and $p$ some point near $\bar{p}$. By non-degeneracy $|\nabla \operatorname{Sc}(p)| \geq c \operatorname{dist}(p, \bar{p})$ and Theorem 1.3 implies

$$
c \operatorname{dist}(p, \bar{p}) \leq C R(\Sigma)
$$

From this it follows that for any neighborhood $U$ of the critical points of Sc there is $a_{0}>0$ such that if $|\Sigma|<a_{0}$ then $\Sigma \subset U$. However, it is not clear that a critical point of Sc lies in the region enclosed by $\Sigma$. This can be derived as a consequence of Theorem 1.4. The assertions of the following corollaries are also included in the paper [4], as step 2 in the proof of their Theorem 1.1.

Corollary 1.5. Let $(M, g)$ be a compact three dimensional Riemannian manifold with $C_{B}$ bounded geometry. Let

$$
Z:=\{x \in M \mid \nabla \mathrm{Sc}(x)=0\}
$$

and assume that the Hessian $\operatorname{Hess} \operatorname{Sc}(x)$ is non-degenerate for every $x \in Z$.
Then there exists an $a_{0} \in(0, \infty)$ depending only on $(M, g)$ such that for every surface $\Sigma$ that satisfies the Euler-Lagrange equation (1) for some $\lambda$, with $|\Sigma| \leq a_{0}$ and $\mathcal{W}(\Sigma) \leq 4 \pi+a_{0}$ the region enclosed by $\Sigma$ intersects $Z$ in a single point.

Note that for the $\Sigma_{a}^{\min }$ it is automatic that $\mathcal{W}(\Sigma) \leq 4 \pi+O(a)$ by comparison with geodesic spheres. Hence, Theorem 1.4 applies in particular to these surfaces and we can be more precise:

Corollary 1.6. Let $(M, g)$ be a compact three dimensional Riemannian manifold. Let

$$
Z^{\max }:=\left\{x \in M \mid \mathrm{Sc}(x)=\max _{M} \mathrm{Sc}\right\}
$$

and assume that the Hessian $\operatorname{Hess} \mathrm{Sc}(x)$ is non-degenerate for every $x \in$ $Z^{\max }$.

Then there exists $a_{0} \in(0, \infty)$ with the following property. For every $a \in$ $\left(0, a_{0}\right)$ the surface $\Sigma_{a}^{\text {min }}$ from Theorem 1.1 is such that it encloses a region that intersects $Z^{\max }$ in a single point.

The paper is organized as follows. In section 2 we collect some estimates from the literature and combine them to an $L^{\infty}$-estimate for $\AA$. In section 3 we introduce a geometric center of mass for small surfaces $\Sigma \subset M$. We use this to select the point $p_{0}$ in the position estimate Theorem 1.4. In sections 4 and 5 we compute the top order contributions in the expansion of certain
geometric quantities on a solution $\Sigma$ of (1). Section 6 provides a calculation of geometric identities necessary in section 7. Theorem 1.4 follows from a slightly more precise version, Theorem 7.1, is carried out in section 7 . Finally, section 8 contains the proof of Corollary 1.5 .

## 2. Preliminaries

Recall the Gauss equation relating the scalar curvature ${ }^{\Sigma}$ Sc of $\gamma$ and the scalar curvature Sc of $g$ :

$$
{ }^{\Sigma} \mathrm{Sc}=\mathrm{Sc}-2 \operatorname{Ric}(\nu, \nu)+\frac{1}{2} H^{2}-|\AA|^{2} .
$$

Denote the genus of $\Sigma$ by $q(\Sigma)$. Integrating the Gauss equation and using Gauss-Bonnet yields that

$$
\begin{equation*}
\mathcal{W}(\Sigma)=4 \pi(1-q(\Sigma))+\frac{1}{2} \mathcal{U}(\Sigma)+\mathcal{V}(\Sigma) \tag{2}
\end{equation*}
$$

where

$$
\begin{equation*}
\mathcal{U}(\Sigma)=\int_{\Sigma}|\AA|^{2} \mathrm{~d} \mu \quad \text { and } \quad \mathcal{V}(\Sigma)=\int_{\Sigma} \operatorname{Ric}(\nu, \nu)-\frac{1}{2} \operatorname{Sc} \mathrm{~d} \mu \tag{3}
\end{equation*}
$$

Equation (2) implies that for bounded area $|\Sigma|$ and bounded ambient curvature $|\operatorname{Ric}|+|\mathrm{Sc}| \leq C_{B}$ a bound for $\mathcal{U}$ is equivalent to bounding $\mathcal{W}$, regardless of the topology of $\Sigma$ :

$$
\begin{equation*}
\mathcal{W}(\Sigma) \leq 4 \pi+\frac{1}{2} \mathcal{U}(\Sigma)+C_{B}|\Sigma| \tag{4}
\end{equation*}
$$

A similar bound holds in the other direction for surfaces $\Sigma$ with bounded genus $q(\Sigma) \leq q_{0}$ :

$$
\mathcal{U}(\Sigma) \leq \mathcal{W}(\Sigma)+4 \pi\left(q_{0}-1\right)+C_{B}|\Sigma|
$$

For the rest of the paper we will use these estimates for spherical surfaces. As a consequence, an a priori bound on $\mathcal{W}($ or $\mathcal{U})$ and on $|\Sigma|$ will yields an a priori bound for the $L^{2}$-norm of the second fundamental form $\|A\|_{L^{2}(\Sigma)}^{2}=$ $\mathcal{U}(\Sigma)+2 \mathcal{W}(\Sigma)$.

### 2.1. A priori estimates for small surfaces of Willmore type

Here we quote some estimates from the papers [6, 7]. They require uniform bounds on the geometry of $(M, g)$ in the following sense.

Definition 2.1 (cf. Definition 2.1 in [7]). Let $(M, g)$ be a complete Riemannian manifold and $C_{B} \in(0, \infty)$. We say that $(M, g)$ has $C_{B}$-bounded geometry if for every $p \in M$ we have $\operatorname{inj}(M, g, p) \geq C_{B}^{-1}$ and $|\operatorname{Rm}(p)|+$ $|\nabla \operatorname{Rm}(p)| \leq C_{B}$.

To phrase the estimates quoted below in a geometric way, we use the area radius of a surface defined as

$$
R(\Sigma)=\sqrt{\frac{|\Sigma|}{4 \pi}}
$$

The following lemma that if we assume small enough area and bounded Willmore energy, then the area radius of a surface is comparable to its diameter.

Lemma 2.2. Let $(M, g)$ be of $C_{B}$-bounded geometry as in Definition 2.1 and $E_{0} \in(0, \infty)$. Then there exist constants $a_{0} \in(0, \infty)$ and $C$ depending only on $C_{B}$ and $E_{0}$ with the following property: If $\Sigma \subset M$ is a smooth closed hypersurface with $\mathcal{W}(\Sigma) \leq E_{0}$ and $|\Sigma| \leq a_{0}$ then

$$
C^{-1} \operatorname{diam}(\Sigma) \leq R(\Sigma) \leq C \operatorname{diam}(\Sigma)
$$

Proof. The right estimate follows from Lemma 2.2 in [6] whereas the left estimate is a direct consequence of Lemma 2.5 in [7].

Theorem 2.3 (cf. Theorem 5.4 from [7]). Let $(M, g)$ be of $C_{B}$-bounded geometry as in Definition 2.1. Then there exist constants $a_{0} \in(0, \infty)$ and $C$ depending only on $C_{B}$ such that for every surface $\Sigma$ satisfying (1), with $|\Sigma| \leq a_{0}$ and $\mathcal{W}(\Sigma) \leq 4 \pi+a_{0}$ we have the estimate

$$
\int_{\Sigma}\left|\nabla^{2} H\right|^{2}+H^{2}|\nabla A|^{2}+H^{4}|\AA|^{2} \mathrm{~d} \mu \leq C
$$

Lemma 2.4 (cf. Corollary 5.5 from [7]). Under the assumptions of theorem 2.3 we have that

$$
\|\AA\|_{L^{2}(\Sigma)} \leq C R(\Sigma)^{2} \quad \text { and } \quad\left\|H-\frac{2}{R(\Sigma)}\right\|_{L^{\infty}(\Sigma)} \leq C R(\Sigma)
$$

This lemma shows that for small enough area $|\Sigma|$ the mean curvature $H>0$ and thus the estimates from [6] hold under the assumptions of Theorem 2.3. It also follows that $\|A\|_{L^{2}(\Sigma)} \leq C$.

For $r \in(0, \infty)$, denote the intrinsic ball centered at $p \in M$ by

$$
\mathcal{B}_{r}(p)=\left\{q \in M \mid \operatorname{dist}_{g}(p, q) \leq r\right\}
$$

Proposition 2.5 (cf. Corollary 3.6 from [6]). Let ( $M, g$ ) and $\Sigma$ be as in Theorem 2.3. Let $p_{0} \in M$ be a point with $\Sigma \subset \mathcal{B}_{2 \operatorname{diam}(\Sigma)}\left(p_{0}\right)$, then

$$
\left|\lambda+\frac{1}{3} \operatorname{Sc}\left(p_{0}\right)\right| \leq C R(\Sigma)
$$

### 2.2. General inequalities

The Bochner identity for surfaces can be stated as follows.
Lemma 2.6. For all functions $f \in C^{\infty}(\Sigma)$ we have that

$$
\int_{\Sigma} 2\left|\nabla^{2} f\right|^{2}+\frac{1}{2} H^{2}|\nabla f|^{2} \mathrm{~d} \mu \leq \int_{\Sigma} 2|\Delta f|^{2}+\left(2 \operatorname{Ric}(\nu, \nu)-\mathrm{Sc}+|\AA|^{2}\right)|\nabla f|^{2} \mathrm{~d} \mu
$$

Proof. The Bochner identity states that

$$
\int_{\Sigma}\left|\nabla^{2} f\right|^{2} \mathrm{~d} \mu=\int_{\Sigma}(\Delta f)^{2}-{ }^{\Sigma} \operatorname{Rc}(\nabla f, \nabla f) \mathrm{d} \mu
$$

Since $\Sigma$ is a surface, its Ricci curvature satisfies ${ }^{\Sigma} \mathrm{Rc}={ }^{1}{ }^{\Sigma} \mathrm{Sc} \gamma$ and the scalar curvature ${ }^{\Sigma}$ Sc of $\Sigma$ can be expressed via the the Gauss equation

$$
{ }^{\Sigma} \mathrm{Sc}=\mathrm{Sc}-\operatorname{Ric}(\nu, \nu)+\frac{1}{2} H^{2}-|\AA|^{2} .
$$

Let $C$ and $a_{0}$ be the constants from lemma 2.2 . Hence, if $\Sigma$ is such that $|\Sigma| \leq$ $a_{0}$ and $p \in M$ is some point such that there exists $x \in \Sigma$ with $\operatorname{dist}(p, x) \leq$ $R(\Sigma)$, then $\Sigma \subset \mathcal{B}_{(C+1) R(\Sigma)}(p)$. Hence, there exists $a_{0}^{\prime} \in\left(0, a_{0}\right)$ such that if $|\Sigma| \leq a_{0}^{\prime}$ and $\rho=(C+1) R(\Sigma)$ then $\rho \leq \operatorname{inj}(M, g, p)$. In this case there are normal coordinates $x: \mathcal{B}_{\rho}(p) \rightarrow B_{\rho}(0) \subset \mathbf{R}^{3}$. The metric $\left(x^{-1}\right)^{*} g$ on $B_{\rho}(0)$ has the expansion $\left(x^{-1}\right)^{*} g=\delta+h$ with

$$
\sup \left(|x|^{-2}|h|+|x|^{-1}|\partial h|+\left|\partial^{2} h\right|\right) \leq h_{0}
$$

Here $h_{0}$ is a constant depending only on $C_{B}$ and $\partial$ denotes partial derivatives in the coordinate system given by $x$. In particular, we can apply the following two estimates on surfaces $\Sigma$ as in section 2.1 with uniform constants, that is constants independent of $\Sigma$, provided $|\Sigma| \leq a_{0}$ for some constant $a_{3} \in(0, \infty)$ depending only on $C_{B}$.

Lemma 2.7. Let $g=g^{E}+h$ on $B_{\rho}$ and $C_{0}$ be given. Then there exists $\rho_{0} \in(0, \rho)$ and a constant $C$ depending only on $\rho, h_{0}$ and $C_{0}$ such that for all surfaces $\Sigma \subset B_{\rho_{0}}$ with $\|A\|_{L^{2}(\Sigma)} \leq C_{0}$ and all $f \in C^{\infty}(\Sigma)$ we have

$$
\left(\int_{\Sigma} f^{2} \mathrm{~d} \mu\right)^{1 / 2} \leq C \int_{\Sigma}|\nabla f|+|H f| \mathrm{d} \mu
$$

In addition, for surfaces as in Theorem 2.3 we have a Poincaré inequality of the following form:

Lemma 2.8. Let $(M, g)$ be of $C_{B}$-bounded geometry as in Definition 2.1. Then there exist constants $a_{0} \in(0, \infty), \varepsilon \in(0, \infty)$ and $C$ depending only on $C_{B}$ with the following property. If $\Sigma \subset M$ is a smooth closed hypersurface with $\mathcal{U}(\Sigma) \leq \varepsilon$ and $|\Sigma| \leq a_{0}$ then for all $f \in C^{\infty}(\Sigma)$ we have

$$
\int_{\Sigma}|f-\bar{f}|^{2} \mathrm{~d} \mu \leq C|\Sigma| \int_{\Sigma}|\nabla f|^{2} \mathrm{~d} \mu
$$

Here

$$
\bar{f}=|\Sigma|^{-1} \int_{\Sigma} f \mathrm{~d} \mu
$$

is the mean value of $f$.
Proof. This holds without assuming an upper bound for the area of $\Sigma$ if $(M, g)$ is Euclidean space in view of the eigenvalue estimates of DeLellis and Müller [3, Corollary 1.3] for nearly umbilical surfaces.

For general $(M, g)$ we first use inequality (4) to bound $\mathcal{W}(\Sigma)$ in terms of $\mathcal{U}(\Sigma)$ and $|\Sigma|$. Then by Lemma 2.2 we infer that small area $|\Sigma|$ implies small diameter. Using normal coordinates covering $\Sigma$ implies that we are in a nearly Euclidean setting. It is then straight forward to deduce the desired Poincaré inequality on $\Sigma$ with respect to the metric induced by $g$ from the one with respect to the Euclidean metric in the normal coordinate system.

The following estimate follows from the Michael-Simon-Sobolev inequality from Lemma 2.7 and can be proved exactly as [5, Lemma 2.8]. This form appears in [6, Lemma 3.7].

Lemma 2.9. Assume that the metric $g=g^{E}+h$ on $B_{\rho}$ is given. Then there exist $\rho_{0} \in(0, \rho)$ and a constant $C<\infty$ such that for all surfaces $\Sigma \subset$
$B_{r}$ with $r \in\left(0, \rho_{0}\right)$ and for all smooth forms $\phi$ on $\Sigma$ we have

$$
\|\phi\|_{L^{\infty}(\Sigma)}^{4} \leq C\|\phi\|_{L^{2}(\Sigma)}^{2} \int_{\Sigma}\left|\nabla^{2} \phi\right|^{2}+|H|^{4}|\phi|^{2} \mathrm{~d} \mu
$$

### 2.3. A $L^{\infty}$ estimate for $\AA$

For the later exposition we need two more estimates not present in [6, 7].

Lemma 2.10. Let $(M, g)$ be as in definition 2.1. Then there exist constants $a_{0} \in(0, \infty)$ and $C \in(0, \infty)$ depending only on $C_{B}$ such that if $\Sigma$ satisfies (1) for some $\lambda \in \mathbf{R}$ with $|\Sigma| \leq a_{0}$ and $\mathcal{W}(\Sigma) \leq 4 \pi+a_{0}$ then

$$
\|\AA\|_{L^{\infty}(\Sigma)} \leq C R(\Sigma)
$$

and

$$
\left\|H^{-1}-R(\Sigma) / 2\right\|_{L^{\infty}(\Sigma)} \leq C R(\Sigma)^{2}
$$

Proof. We assume that $a_{0}$ is so small that the estimates from theorem 2.3 apply. Then the second estimate is an immediate consequence from the second estimate in 2.4. To show the first estimate, proceed as in the proof of 8 , Lemma 15]. In view of the Bochner identity it suffices to estimate $\Delta \AA$ in the $L^{2}$-norm. The $L^{\infty}$ estimate then follows from lemma 2.9. To derive this estimate, recall the Simons identity [13] in the form of [8, eq. (8)]

$$
\begin{align*}
\Delta \AA_{i j}= & \left(\nabla^{2} H\right)_{i j}^{\circ}+H \AA_{i}^{k} \AA_{k j}+\frac{1}{2} H^{2} \AA_{i j}-|\AA|^{2} \AA_{i j}-\frac{1}{2} H|\AA|^{2} \gamma_{i j} \\
& +\AA_{j}^{k} \gamma_{l m} \operatorname{Rm}_{l i k m}+\AA^{\circ k l} \operatorname{Rm}_{i k j l}+2 \nabla_{i} \omega_{j}-\operatorname{div} \omega \gamma_{i j} \tag{5}
\end{align*}
$$

Here $\omega=\operatorname{Ric}(\nu, \cdot)^{T}$ denotes the tangential 1-form obtained from projecting the 1 -from $\operatorname{Ric}(\nu, \cdot)$ to the tangent space of $\Sigma$. From the calculation in section 4.2 we get that

$$
|\nabla \omega| \leq|\nabla \operatorname{Ric}|+|A| \mid \text { Ric } \mid .
$$

This yields

$$
\|\nabla \omega\|_{L^{2}(\Sigma)}^{2} \leq C|\Sigma|+\int_{\Sigma}|A|^{2} \mathrm{~d} \mu \leq C
$$

Together with equation (5) this gives

$$
\begin{aligned}
&\|\Delta \AA\|_{L^{2}(\Sigma)} \leq c\left(\left\|\nabla^{2} H\right\|_{L^{2}(\Sigma)}+\left\|H^{2} \AA\right\|_{L^{2}(\Sigma)}+\|\AA\|_{L^{6}}^{2}\right. \\
&\left.+\|A\|_{L^{2}}\|\operatorname{Rm}\|_{L^{\infty}(\Sigma)}+\|\nabla \operatorname{Ric}\|_{L^{2}(\Sigma)}\right) \\
& \leq C+c\|\AA\|_{L^{\infty}(\Sigma)}^{2}\|\AA\|_{L^{2}(\Sigma)}
\end{aligned}
$$

Here $c$ denotes a purely numerical constant, and we used the estimates from section 2.1 in the second step.

From the Bochner identity 2.6 (more precisely a variant for two tensors) we obtain that

$$
\begin{aligned}
& \left\|\nabla^{2} \AA\right\|_{L^{2}(\Sigma)}+\|H \nabla \AA\|_{L^{2}(\Sigma)} \\
& \quad \leq c\|\Delta \AA\|_{L^{2}(\Sigma)}+C\left(1+\|\AA\|_{L^{\infty}(\Sigma)}\right)\|\nabla \AA\|_{L^{2}(\Sigma)} \\
& \quad \leq C+c\|\AA\|_{L^{\infty}(\Sigma)}^{2}\|\AA\|_{L^{2}(\Sigma)}+C\left(1+\|\AA\|_{L^{\infty}(\Sigma)}\right)\|\nabla \AA\|_{L^{2}(\Sigma)}
\end{aligned}
$$

Note that the last term on the right hand side can be absorbed to the left, if $R$ is small enough.

This yields in view of lemma 2.9 that

$$
\begin{aligned}
\|\AA\|_{L^{\infty}(\Sigma)}^{4} & \leq C\|\AA\|_{L^{2}(\Sigma)}^{2}\left(\left\|\nabla^{2} \AA\right\|_{L^{2}(\Sigma)}^{2}+\left\|H^{2} \AA\right\|_{L^{2}(\Sigma)}\right) \\
& \leq C R(\Sigma)^{4}\left(C+C\|\AA\|_{L^{\infty}(\Sigma)}^{4} R(\Sigma)^{2}\right)
\end{aligned}
$$

If $R(\Sigma)$ is small enough, this gives

$$
\|\AA\|_{L^{\infty}(\Sigma)} \leq C R(\Sigma)
$$

Note that by choosing $a_{0}$ small enough, we can ensure that $R(\Sigma)$ is so small, that the above steps apply to $\Sigma$ as in the assumption.

### 2.4. Approximately spherical surfaces

In this section we discuss the approximation of a given surface $\Sigma \subset \mathbf{R}^{3}$ by spheres. The main tool here are the estimates from DeLellis and Müller [2, 3]. We quote their estimates in the form needed here from [6, Theorem 2.4]. These results are purely Euclidean. To distinguish geometric quantities computed with respect to the Euclidean metric we use the superscript ${ }^{E}$.

Theorem 2.11. There exists a universal constant $C$ with the following properties. Assume that $\Sigma \subset \mathbf{R}^{3}$ is a surface with $\left\|\AA^{E}\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)}^{2}<$
$8 \pi$. Let $R^{E}:=\sqrt{|\Sigma|^{E} / 4 \pi}$ be the Euclidean area radius of $\Sigma$ and $a^{E}:=$ $|\Sigma|_{E}^{-1} \int_{\Sigma} x \mathrm{~d} \mu^{E}$ be the Euclidean center of gravity. Then there exists a conformal map $F: S:=S_{R^{E}}\left(a^{E}\right) \rightarrow \Sigma \subset \mathbf{R}^{3}$ with the following properties. Let $\gamma^{S}$ be the standard metric on $S, N$ the Euclidean normal vector field and $\phi$ the conformal factor, that is $F^{*} \gamma^{E}=\phi^{2} \gamma^{S}$. Then the following estimates hold

$$
\begin{aligned}
\left\|H^{E}-2 / R^{E}\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)} & \leq C\left\|\AA^{E}\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)} \\
\left\|F-\operatorname{id}_{S}\right\|_{L^{\infty}(S)} & \leq C R^{E}\left\|\AA^{E}\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)} \\
\left\|\phi^{2}-1\right\|_{L^{\infty}(S)} & \leq C\left\|\AA^{E}\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)} \\
\left\|N-\nu^{E} \circ F\right\|_{L^{2}(S)} & \leq C R^{E}\left\|\AA^{E}\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)}
\end{aligned}
$$

These estimates can be applied in our situation by choosing appropriate normal coordinates near a small surfaces as described in section 2.2 and compare geometric quantities in the given metric to the Euclidean background. In particular we have:

Lemma 2.12 (cf. [6, Lemma 2.5]). Let $g=g^{E}+h$ on $B_{\rho}$ be given. Then there exists $0<\rho_{0}<\rho$ and a constant $C$ depending only on $\rho$ and $h_{0}$ such that for all surfaces $\Sigma \subset B_{r}$ with $r<\rho_{0}$ we have

$$
\left\|\AA^{E}\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)}^{2} \leq C\|\AA\|_{L^{2}(\Sigma, \gamma)}^{2}+C \rho^{4}\|H\|_{L^{2}(\Sigma, \gamma)}^{2}
$$

## 3. A geometric center of mass

The calculations in section 7 require that the normal coordinates in which we look at our surfaces $\Sigma$ are well adapted to $\Sigma$. In this section we propose one way to assign a geometric center of mass to our surfaces. Centering the normal coordinates there gives good control on the center of mass of the image of the surface in the coordinate picture.

Let $(M, g)$ be a Riemannian manifold and $\Sigma \subset M$ a closed smooth hypersurface with extrinsic diameter $d=\operatorname{diam}(\Sigma)=\max \{\operatorname{dist}(x, y) \mid x, y \in \Sigma\}$ where dist denotes the distance function in $(M, g)$. Assume that $2 d<$ $\operatorname{inj}(M, g)$. For $p \in M$ let $d_{p}(x):=\operatorname{dist}(p, x)$ and set

$$
w(p):=\int_{\Sigma} d_{p}(x)^{2} \mathrm{~d} \mu
$$

Then $w$ is a smooth, positive, proper function on $M$ which attains its global infimum on the compact set

$$
K:=\{p \in M \mid \operatorname{dist}(p, \Sigma) \leq d\}
$$

This follows from comparing values of $w$ outside of $K$ with $w(p)$ for some $p \in \Sigma$. Let $p_{0} \in K$ be a point where $w$ attains its minimum. Since $p_{0} \in K$ we have that $\Sigma \subset \mathcal{B}_{2 d}\left(p_{0}\right)$ and since $2 d<\operatorname{inj}(M, g)$ we find that $\Sigma$ is completely contained in a normal coordinate neighborhood centered at $p_{0}$. Let $\psi: \mathcal{B}_{\rho}\left(p_{0}\right) \rightarrow B_{\rho}(0) \subset \mathbf{R}^{n}$ be such normal coordinates where $\rho>2 d$ denotes the injectivity radius on $(M, g)$ at $p_{0}$. Let $x \in B_{\rho}(0)$ and $p=\psi^{-1}(x)$. Then

$$
\tilde{w}(x):=w(p)=\int_{\psi(\Sigma)} \operatorname{dist}_{g}(x, y)^{2} \mathrm{~d} \mu_{g}(y)
$$

where $\operatorname{dist}_{g}$ is the distance function induced by the pull-back metric $\left(\psi^{-1}\right)^{*} g$ to $B_{\rho}(0)$ and $\mathrm{d} \mu_{g}$ denotes the induced surface measure.

Since $w$ is critical at $p_{0}$ also $\tilde{w}$ is critical at 0 and we compute

$$
0=\frac{\partial}{\partial x^{\alpha}} \tilde{w}(0)=2 \int_{\psi(\Sigma)} y^{\alpha} \mathrm{d} \mu_{g}
$$

since in normal coordinates $\operatorname{dist}_{g}(x, y)^{2}=|x-y|^{2}+O\left(|x|^{2}\right)$.
We can also change the surface measure to the Euclidean one, recording the error term:

$$
\left|\int_{\psi(\Sigma)} y^{\alpha} \mathrm{d} \mu_{g}-\int_{\psi(\Sigma)} y^{\alpha} \mathrm{d} \mu^{E}\right| \leq C d^{3}|\Sigma| .
$$

Here and in the following we use $y$ to refer to the position vector on $\psi(\Sigma)$.
Summarizing, we arrive at the following:
Lemma 3.1. Let $(M, g)$ be of $C_{B}$-bounded geometry. Then there exists a constant $C$ depending only on $C_{B}$ with the following property: For every closed smooth hypersurface $\Sigma \subset M$ with extrinsic diameter $d=\operatorname{diam}(\Sigma)=$ $\max \{\operatorname{dist}(x, y) \mid x, y \in \Sigma\}<\frac{1}{2} \operatorname{inj}(M, g)$ there exists a point $p_{0} \in M$ such that $\operatorname{dist}\left(p_{0}, \Sigma\right) \leq d$ such that in normal coordinates centered at $p_{0}$ we have that

$$
\int_{\psi(\Sigma)} y^{\alpha} \mathrm{d} \mu_{g}=0 \quad \text { and } \quad\left|\int_{\psi(\Sigma)} y^{\alpha} \mathrm{d} \mu^{E}\right| \leq C d^{3}|\Sigma|
$$

Combined with theorem 2.11 and lemma 2.12 we obtain the following estimate in the case where $\Sigma$ is a surface in a 3 -dimensional manifold:

Lemma 3.2. Let $(M, g)$ be three dimensional and of $C_{B}$-bounded geometry. Then there exist constants $C$ and $a_{0} \in(0, \infty)$ depending only on $C_{B}$ with the following property: For every closed smooth surface $\Sigma \subset M$ with $|\Sigma| \leq a_{0}$ and $\mathcal{U}(\Sigma) \leq a_{0}$ there exists a point $p_{0} \in M$, normal coordinates $\psi: \mathcal{B}_{\rho}\left(p_{0}\right) \rightarrow$ $B_{\rho}(0) \subset \mathbf{R}^{3}$ and in these coordinates we have that

$$
\begin{equation*}
\left\|\frac{y}{R}-\nu\right\|_{L^{2}(\Sigma)} \leq C\left(R^{3}+R\|\AA\|_{L^{2}(\Sigma)}\right) \tag{6}
\end{equation*}
$$

and

$$
\begin{equation*}
\left\|\operatorname{dist}\left(p_{0}, \cdot\right)-R\right\|_{L^{\infty}(\Sigma)} \leq C\left(R^{3}+R\|\AA\|_{L^{2}(\Sigma)}\right) \tag{7}
\end{equation*}
$$

Here $R=R(\Sigma)$ denotes the area radius of $\Sigma$.
Proof. We choose $a_{0} \in(0,1]$ in a moment. By (4) this gives the a priori bound $\mathcal{W}(\Sigma) \leq 4 \pi+\frac{1}{2}+C_{B}$. In view of the diameter bound from Lemma 2.2 we can choose $a_{0} \in(0,1]$ so small that Lemma 3.1 holds for $\Sigma$ as in the assumption. Let $\psi: \mathcal{B}_{\rho}\left(p_{0}\right) \rightarrow B_{\rho}(0)$ denote the coordinates from there. In view of the diameter estimate the quantities $d=\operatorname{diam}(\Sigma)$ from Lemma 3.1 and $R$ are comparable. Hence also $\max _{p \in \Sigma} \operatorname{dist}\left(p, p_{0}\right) \leq C R$ so that the estimate from Lemma 2.12 can be rephrased as

$$
\begin{equation*}
\left\|\AA^{E}\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)}^{2} \leq C\left(\mathcal{U}(\Sigma)+R^{4}\right) \tag{8}
\end{equation*}
$$

where we also used that the Willmore functional is a priori bounded.
To prove (6), it is thus sufficient to prove the Euclidean inequality

$$
\left\|\frac{y}{R}-\nu^{E}\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)} \leq C\left(R^{3}+R\left\|\AA^{E}\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)}\right)
$$

since in the previous coordinates we have that $\nu-\nu^{E}=O\left(R^{2}\right)$ and due to equation (8). Compute

$$
\nabla^{E} \frac{y}{R}=R^{-1} \mathrm{Id} \quad \text { and } \quad \nabla^{E} \nu^{E}=\frac{H^{E}}{2} \mathrm{Id}+\AA^{E}
$$

Here we denote the tangential derivative along $\Sigma$ by $\nabla^{E}$, Id denotes the identity endomorphism field in the tangent bundle on $\Sigma$ and we slightly abuse notation by not distinguishing $\AA^{E}$ from its associated endomorphism.

This gives the estimate

$$
\begin{aligned}
\left\|\nabla\left(\frac{y}{R}-\nu^{E}\right)\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)} & \leq \frac{1}{2}\left\|\frac{2}{R}-H^{E}\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)}+\left\|\AA^{E}\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)} \\
& \leq C\left\|\AA^{E}\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)}
\end{aligned}
$$

The last inequality follows from Theorem 2.11 if $a_{0} \in(0,1]$ is chosen so small that equation (8) implies $\left\|\AA^{E}\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)}^{2} \leq 6 \pi$.

By choosing $a_{0} \in(0,1]$ even smaller, we can ensure that the Poincaré inequality from Theorem 2.8 holds. This gives

$$
\left\|\frac{y}{R}-\nu^{E}-m\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)} \leq C R\left\|\AA^{E}\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)}
$$

where

$$
m=|\Sigma|^{-1} \int_{\psi(\Sigma)}\left(\frac{y}{R}-\nu^{E}\right) \mathrm{d} \mu^{E}=|\Sigma|^{-1} \int_{\psi(\Sigma)} \frac{y}{R} \mathrm{~d} \mu^{E}
$$

By Lemma 3.1 we have $|m| \leq C R^{2}$ that is $\|m\|_{L^{2}(\Sigma)} \leq C R^{3}$ and thus the first of the claimed estimate follows.

To show equation (7) observe that the Euclidean center of gravity $a^{\Sigma}$ of $\psi(\Sigma)$ satisfies

$$
\left|a^{E}\right|=|\Sigma|_{E}^{-1}\left|\int_{\psi(\Sigma)} y \mathrm{~d} \mu^{E}\right| \leq C R^{3}
$$

Parameterizing $\psi(\Sigma)$ with a map $F: S_{R^{E}}\left(a^{E}\right) \rightarrow \psi(\Sigma)$ as in Theorem 2.11, we get that for every $y \in S_{R^{E}}\left(a^{E}\right)$

$$
\left|y-a^{E}\right|-\left|a^{E}\right|-|F(y)-y| \leq|F(y)| \leq\left|y-a^{E}\right|+\left|a^{E}\right|+|F(y)-y|
$$

so that

$$
\left||F(y)|-R^{E}\right| \leq\left|a^{E}\right|+|F(y)-y| \leq C R^{3}+C R^{E}\left\|\AA^{E}\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)}
$$

Since we are in normal coordinates around $p_{0}$ we have that for all $x \in \Sigma$ that $\operatorname{dist}\left(x, p_{0}\right)=|\psi(x)|=\left|F\left(F^{-1}(\psi(x))\right)\right|$ and the second claim follows.

Corollary 3.3. Let $(M, g)$ and $\Sigma$ satisfy the assumptions of Lemma 3.2 and let $\psi$ be as there. Then for every $k \in \mathbf{N} \cup\{0\}$ there is a constant $C_{k}$
depending only on the constant $C$ in Lemma 3.2 and on $k$ such that

$$
\left|\int_{\psi(\Sigma)} \prod_{l=1}^{2 k+1} \rho_{l} \mathrm{~d} \mu\right| \leq C_{k}\left(R^{3}+R\|\AA\|_{L^{2}(\Sigma)}\right)
$$

Here, for every $l \in\{1, \ldots, 2 k+1\}$ we can choose $\rho_{l}$ freely from the functions $\left\{\nu^{\alpha}, \left.\frac{y^{\alpha}}{R} \right\rvert\, \alpha=1,2,3\right\}$.

Proof. Note that if $k=0$ then the claim directly follows from Lemma 3.1 if $\rho_{1}=\frac{y^{\alpha}}{R}$ and from the fact that $\int_{\psi(\Sigma)}\left(\nu^{E}\right)^{\alpha} \mathrm{d} \mu^{E}=0$ if $\rho_{1}=\nu^{\alpha}$ for some $\alpha=1,2,3$. For brevity, we indicate the proof only in the case $k=1$ below. Also note that it is sufficient to consider the Euclidean setting, that is with $\rho_{l} \in\left\{\left(\nu^{E}\right)^{\alpha}, \left.\frac{y^{\alpha}}{R} \right\rvert\, \alpha=1,2,3\right\}$ and with $\|\AA\|_{L^{2}(\Sigma)}$ replaced by $\left\|\AA^{E}\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)}$ using the same reduction as in the proof of Lemma 3.2.

We proceed in two steps: in the first step, we use Theorem 2.11 to prove the case $\rho_{l} \in\left\{\left(\nu^{E}\right)^{\alpha} \mid \alpha=1,2,3\right\}$, in the second step we use Lemma 3.2 to conclude.

Step 1: Let $\rho_{l}=\left(\nu^{E}\right)^{\alpha_{l}}$ for $l=1,2,3$ and $\alpha_{l} \in\{1,2,3\}$. Let $R^{E}$ and $a^{E}$ as in Theorem 3.2 and denote $S:=S_{R^{E}}\left(a^{E}\right)$. Let $F: \rightarrow \Sigma$ be the parameterization from Theorem 3.2, $N: S \rightarrow S^{2}$ be the normal of $S$ and $\tilde{\rho}_{l}:=\rho_{l} \circ F$. Then we can write

$$
\begin{aligned}
& \int_{\psi(\Sigma)} \rho_{1} \rho_{2} \rho_{3} \mathrm{~d} \mu^{E}=\int_{S} \tilde{\rho}_{1} \tilde{\rho}_{2} \tilde{\rho}_{3} \phi^{2} \mathrm{~d} \mu^{E} \\
& \quad=\int_{S} N^{\alpha_{1}} N^{\alpha_{2}} N^{\alpha_{3}}+N^{\alpha_{1}} N^{\alpha_{2}} N^{\alpha_{3}}\left(\phi^{2}-1\right)+\left(\tilde{\rho}_{1}-N^{\alpha_{1}}\right) N^{\alpha_{2}} N^{\alpha_{3}} \phi^{2} \\
& \quad \quad+\tilde{\rho}_{1}\left(\tilde{\rho}_{2}-N^{\alpha_{2}}\right) N^{\alpha_{3}} \phi^{2}+\tilde{\rho}_{1} \tilde{\rho}_{2}\left(\tilde{\rho}_{3}-N^{\alpha_{3}}\right) \phi^{2} \mathrm{~d} \mu^{E}
\end{aligned}
$$

Since $S$ is a sphere $\int_{S} N^{\alpha_{1}} N^{\alpha_{2}} N^{\alpha_{3}} \mathrm{~d} \mu^{E}=0$ and thus, using Cauchy-Schwarz in the first inequality and Theorem 2.11 in the last inequality we conclude

$$
\begin{aligned}
& \left|\int_{\psi(\Sigma)} \rho_{1} \rho_{2} \rho_{3} \mathrm{~d} \mu^{E}\right| \\
& \quad \leq C R\left\|N-\nu^{E} \circ R\right\|_{L^{2}\left(S, \gamma^{E}\right)}+C R^{2}\left\|\phi^{2}-1\right\|_{L^{\infty}(S)} \leq C R^{2}\left\|\AA^{E}\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)}
\end{aligned}
$$

Note that this implies the claimed inequality.
Step 2: Assume that $\rho_{l}=\left(\nu^{E}\right)^{\alpha_{l}}$ or $\rho_{l}=\frac{y^{\alpha_{l}}}{R}$ for $l=1,2,3$. We can use a telescope sum as above and Cauchy-Schwarz to estimate

$$
\left|\int_{\psi(\Sigma)} \rho_{1} \rho_{2} \rho_{3}-\left(\nu^{E}\right)^{\alpha_{1}}\left(\nu^{E}\right)^{\alpha_{2}}\left(\nu^{E}\right)^{\alpha_{3}} \mathrm{~d} \mu\right| \leq C R \sum_{l=1}^{3}\left\|\rho_{l}-\left(\nu^{E}\right)^{\alpha_{l}}\right\|_{L^{2}\left(\Sigma, \gamma^{E}\right)}
$$

Note that the terms in the sum on the right either vanish or can be bounded using Lemma 3.2. We thus arrive at the claimed inequality.

Note that products of an even number of factors can be treated in a similar fashion as above and equal the respective integrals on a centered round sphere up to the same error term as above.

## 4. Geometric identities

Throughout this section we assume that $(M, g)$ has $C_{B}$-bounded geometry and that $\Sigma \subset M$ is a closed, immersed, smooth surface such that

1) $\Sigma$ satisfies equation (1).
2) $|\Sigma| \leq a_{0}$ and $\mathcal{W}(\Sigma) \leq 4 \pi+a_{0}$.

Here we assume that $a_{0}$ is so small that the estimates from Theorem 2.3 and Lemmas 2.4 and 2.10 hold.

To shorten the exposition, we augment the big- $O$ notation as follows. If $f$ is some quantity defined on a surface $\Sigma$ as above, we say $f=O_{L^{p}}\left(R^{k}\right)$ if

$$
\int_{\Sigma} f^{p} \mathrm{~d} \mu \leq C R^{p k+2}
$$

where $R=R(\Sigma)$ refers to the area radius of $\Sigma$. We also use this for $p=\infty$, that is $f=O_{L^{\infty}}\left(R^{k}\right)$ denotes

$$
\|f\|_{L^{\infty}(\Sigma)} \leq C R^{k} .
$$

Using this notation, the a priori estimates from section 2 can be stated as follows:

$$
\AA=O_{L^{\infty}}(R), \quad \nabla A=O_{L^{2}}(1), \quad \text { and } \quad \nabla^{2} H=O_{L^{2}}\left(R^{-1}\right) .
$$

Lemmas 2.4 and 2.10 imply that $H=O_{L^{\infty}}\left(R^{-1}\right)$ and $H^{-1}=O_{L^{\infty}}(R)$.
The following computations are done in abstract index notation, where Latin indices $i, j, k, \ldots \in\{1,2\}$ refer to an local orthonormal frame $\left\{e_{1}, e_{2}\right\}$ on $\Sigma$ and $\nu$ denotes a choice of normal to $\Sigma \subset M$ so that $A_{i j}:=A\left(e_{i}, e_{j}\right)=$ $g\left({ }^{M} \nabla_{e_{i}} \nu, e_{j}\right)$ where ${ }^{M} \nabla$ denotes the Levi-Civita connection on $(M, g)$. Tangential derivatives to $\Sigma$ are denoted by $\nabla$.

### 4.1. The Hessian of $\operatorname{Ric}(\nu, \nu)$

We begin by calculating the gradient

$$
\begin{align*}
\nabla_{i} \operatorname{Ric}(\nu, \nu) & =\left({ }^{M} \nabla_{e_{i}} \operatorname{Ric}\right)(\nu, \nu)+2 A_{i k} \operatorname{Ric}\left(e_{k}, \nu\right) \\
& =\left({ }^{M} \nabla_{e_{i}} \operatorname{Ric}\right)(\nu, \nu)+H \operatorname{Ric}\left(e_{i}, \nu\right)+2 \AA_{i k} \operatorname{Ric}\left(e_{k}, \nu\right)  \tag{9}\\
& =H \omega_{i}+O_{L^{\infty}}(1) .
\end{align*}
$$

As before $\omega=\operatorname{Ric}(\nu, \cdot)^{T}$ denotes the tangential projection of the 1-form $\operatorname{Ric}(\nu, \cdot)$ to $\Sigma$. In the second step we used the splitting

$$
A_{i j}=\AA_{i j}+\frac{1}{2} H \gamma_{i j}
$$

Differentiating further yields

$$
\begin{align*}
\nabla_{i, j}^{2} \operatorname{Ric}(\nu, \nu)= & \left({ }^{M} \nabla_{i, j}^{2} \operatorname{Ric}\right)(\nu, \nu)-\left({ }^{M} \nabla_{\nu} \operatorname{Ric}\right)(\nu, \nu) A_{i j} \\
& +2\left({ }^{M} \nabla_{e_{i}} \operatorname{Ric}\right)\left(e_{k}, \nu\right) A_{j k}+2\left({ }^{M} \nabla_{e_{j}} \operatorname{Ric}\right)\left(e_{k}, \nu\right) A_{i k} \\
& -2 \operatorname{Ric}(\nu, \nu) A_{k j} A_{j}^{k}+2 \operatorname{Ric}\left(e_{k}, e_{l}\right) A_{i}^{k} A_{j}^{l}  \tag{10}\\
& +2 \operatorname{Ric}\left(e_{k}, \nu\right) \nabla_{e_{i}} A_{j}^{k} \\
= & -2 \operatorname{Ric}(\nu, \nu) A_{i}^{k} A_{k j}+2 \operatorname{Ric}\left(e_{k}, e_{l}\right) A_{i}^{k} A_{j}^{l}+O_{L^{2}}\left(R^{-1}\right) \\
= & -\frac{1}{2} H^{2} \operatorname{Ric}(\nu, \nu) \gamma_{i j}+\frac{1}{2} H^{2} T_{i j}+O_{L^{2}}\left(R^{-1}\right)
\end{align*}
$$

Here $T_{i j}=\operatorname{Ric}\left(e_{i}, e_{j}\right)$ denotes the tangential projection of the Ricci-Tensor. The last step uses Lemma 2.10 to discard the terms containing $\AA$ into the error term.

Taking the trace in equation (10) yields that

$$
\begin{equation*}
\Delta \operatorname{Ric}(\nu, \nu)=-\frac{3}{2} H^{2} \operatorname{Ric}(\nu, \nu)+\frac{1}{2} H^{2} \mathrm{Sc}+O_{L^{2}}\left(R^{-1}\right) \tag{11}
\end{equation*}
$$

We thus infer that the trace free part of the Hessian of $\operatorname{Ric}(\nu, \nu)$ is given by

$$
\begin{align*}
\left(\nabla^{2} \operatorname{Ric}(\nu, \nu)\right)_{i j}^{\circ} & =\nabla_{i, j}^{2} \operatorname{Ric}(\nu, \nu)-\frac{1}{2} \Delta \operatorname{Ric}(\nu, \nu) \gamma_{i j} \\
& =\frac{1}{2} H^{2}\left(\frac{1}{2} \operatorname{Ric}(\nu, \nu) \gamma_{i j}-\frac{1}{2} \operatorname{Sc} \gamma_{i j}+T_{i j}\right)+O\left(R^{-1}\right)  \tag{12}\\
& =\frac{1}{2} H^{2} \stackrel{\circ}{T}_{i j}+O_{L^{2}}\left(R^{-1}\right)
\end{align*}
$$

Here we used that

$$
\stackrel{\circ}{T}_{i j}=T_{i j}-\frac{1}{2} \operatorname{tr} T \gamma_{i j}=T_{i j}+\frac{1}{2} \operatorname{Ric}(\nu, \nu) \gamma_{i j}-\frac{1}{2} \operatorname{Sc} \gamma_{i j}
$$

### 4.2. The covariant derivative of $\omega$

In a calculation similar to equation (10), we derive

$$
\begin{align*}
\nabla_{i} \omega_{j} & =\left({ }^{M} \nabla_{i} \operatorname{Ric}\right)\left(\nu, e_{j}\right)+A_{i}^{k} \operatorname{Ric}\left(e_{k}, e_{j}\right)-A_{i j} \operatorname{Ric}(\nu, \nu) \\
& =\frac{1}{2} H\left(T_{i j}-\operatorname{Ric}(\nu, \nu) \gamma_{i j}\right)+O_{L^{\infty}}(1) \tag{13}
\end{align*}
$$

Taking the trace yields

$$
\begin{equation*}
\operatorname{div} \omega=\frac{1}{2} H \mathrm{Sc}-\frac{3}{2} H \operatorname{Ric}(\nu, \nu)+O_{L^{\infty}}(1) . \tag{14}
\end{equation*}
$$

Later we will also use the following combination

$$
\begin{align*}
2 \nabla_{i} \omega_{j}-\operatorname{div} \omega \gamma_{i j} & =H\left(T_{i j}+\frac{1}{2} \operatorname{Ric}(\nu, \nu) \gamma_{i j}-\frac{1}{2} \operatorname{Sc} \gamma_{i j}\right)+O_{L^{\infty}}(1) \\
& =H \stackrel{\circ}{T}_{i j}+O_{L^{\infty}}(1) . \tag{15}
\end{align*}
$$

### 4.3. The Laplacian of $\stackrel{\circ}{T}$

First calculate the Hessian of $T$. Neglecting the lower order terms yields

$$
\begin{aligned}
& \nabla_{k, l}^{2} T_{i j}=\frac{1}{4} H^{2}\left(\gamma_{k i} \gamma_{l j} \operatorname{Ric}(\nu, \nu)+\gamma_{k j} \gamma_{l i} \operatorname{Ric}(\nu, \nu)\right. \\
&\left.\quad-\gamma_{k i} \operatorname{Ric}\left(e_{l}, e_{j}\right)-\gamma_{k j} \operatorname{Ric}\left(e_{l}, e_{i}\right)\right)+O_{L^{2}}\left(R^{-1}\right)
\end{aligned}
$$

Taking the trace gives

$$
\Delta T_{i j}=\frac{1}{2} H^{2}\left(\operatorname{Ric}(\nu, \nu) \gamma_{i j}-T_{i j}\right)+O_{L^{2}}\left(R^{-1}\right)
$$

Thus we can calculate further

$$
\Delta \stackrel{\circ}{T}_{i j}=\Delta T_{i j}-\frac{1}{2} \Delta(\operatorname{Sc}-\operatorname{Ric}(\nu, \nu)) \gamma_{i j} .
$$

In view of the fact that

$$
\Delta \mathrm{Sc}={ }^{M} \Delta \mathrm{Sc}-{ }^{M} \nabla_{\nu, \nu}^{2} \mathrm{Sc}+H g\left({ }^{M} \nabla \mathrm{Sc}, \nu\right)=O_{L^{2}}\left(R^{-1}\right)
$$

and the expression for $\Delta \operatorname{Ric}(\nu, \nu)$ in (11), we infer that

$$
\begin{equation*}
\Delta \stackrel{\circ}{T}_{i j}=-\frac{1}{2} H^{2} \stackrel{\circ}{T}_{i j}+O_{L^{2}}\left(R^{-1}\right) \tag{16}
\end{equation*}
$$

## 5. Expansion of the curvature

In this section we consider the crucial geometric quantities on $\Sigma$ as in section 4 and derive the top order deviations from their Euclidean value.

### 5.1. Curvature corrections to $\boldsymbol{H}^{2}$

Combining equations (11) and (11) with the curvature estimates we infer that

$$
\begin{align*}
\Delta & \left(\frac{1}{2} H^{2}-\frac{2}{3} \operatorname{Ric}(\nu, \nu)\right) \\
= & H \Delta H+|\nabla H|^{2}-\frac{2}{3} \Delta \operatorname{Ric}(\nu, \nu) \\
= & -H^{2} \operatorname{Ric}(\nu, \nu)-H^{2} \lambda \\
& -\frac{2}{3}\left(-\frac{3}{2} H^{2} \operatorname{Ric}(\nu, \nu)+\frac{1}{2} H^{2} \mathrm{Sc}\right)+O\left(R^{-1}\right)  \tag{17}\\
= & -H^{2}(\lambda+\mathrm{Sc})+O_{L^{2}}\left(R^{-1}\right) \\
= & O_{L^{2}}\left(R^{-1}\right)
\end{align*}
$$

This identity leads to the following estimate.
Proposition 5.1. Let $(M, g)$ be of $C_{B}$-bounded geometry. Then there exist constants $a_{0} \in(0, \infty)$ and $C$ depending only on $C_{B}$ such that for every surface $\Sigma$ satisfying (1), with $|\Sigma| \leq a_{0}$ and $\mathcal{W}(\Sigma) \leq 4 \pi+a_{0}$ we have the estimate

$$
\left\|\frac{1}{2} H^{2}-8 \pi|\Sigma|^{-1}-\frac{2}{3} \operatorname{Ric}(\nu, \nu)+\frac{5}{9} \operatorname{Sc}(0)\right\|_{L^{\infty}} \leq C R(\Sigma)
$$

Proof. Let $w=\frac{1}{2} H^{2}-\frac{2}{3} \operatorname{Ric}(\nu, \nu)$. The Bochner identity from Lemma 2.6 implies that

$$
\int_{\Sigma} 2\left|\nabla^{2} w\right|^{2}+H^{2}|\nabla w|^{2} \mathrm{~d} \mu \leq \int_{\Sigma}|\Delta w|^{2}+\left(\operatorname{Ric}(\nu, \nu)-\mathrm{Sc}+|\AA|^{2}\right)|\nabla w|^{2} \mathrm{~d} \mu
$$

Note that Ric and Sc are bounded by a constant, that $\|\AA\|_{L^{\infty}} \leq C R(\Sigma)$ by lemma 2.10, and that $H \geq C^{-1} R(\Sigma)^{-1}$ if $a_{0}$ is chosen small enough. If necessary we can decrease $a_{0}$ further so that the gradient term on the right can be absorbed to the left. In view of equation (17), this yields

$$
\left\|\nabla^{2} w\right\|_{L^{2}} \leq C \quad \text { and } \quad\|\nabla w\|_{L^{2}} \leq C R(\Sigma)
$$

Consequently, the Poincaré inequality implies the estimate

$$
\|w-\bar{w}\|_{L^{2}} \leq C R(\Sigma)^{2}
$$

Plugging this into the estimate from lemma 2.9, we infer that

$$
\begin{equation*}
\|w-\bar{w}\|_{L^{\infty}} \leq C R(\Sigma) \tag{18}
\end{equation*}
$$

It remains to calculate $\bar{w}$. To this end recall [6, Theorem 5.1]. This implies that

$$
\left|\int_{\Sigma} \frac{1}{2} H^{2} \mathrm{~d} \mu-8 \pi+\frac{|\Sigma|}{3} \mathrm{Sc}(0)\right| \leq C R(\Sigma)^{3}
$$

From [6, Lemma 3.3] it follows that in addition

$$
\left|\int_{\Sigma} \operatorname{Ric}(\nu, \nu) \mathrm{d} \mu-\frac{|\Sigma|}{3} \operatorname{Sc}(0)\right| \leq C R(\Sigma)^{3}
$$

In combination, this implies that for $\bar{w}=|\Sigma|^{-1} \int_{\Sigma} w \mathrm{~d} \mu$ we have.

$$
\left|\bar{w}-8 \pi+\frac{5}{9} \mathrm{Sc}(0)\right| \leq C R .
$$

In view of (18) this yields the claim.
Remark 5.2. Note that this is not the expansion of $H^{2}$ on geodesic spheres, which can be found in [12, Lemma 2.4] for example. This is due to the fact that geodesic spheres do not satisfy (1) on the order on which we do these calculations. In other words, if a surface satisfies (1), then its shape differs from that of a geodesic sphere in a way visible in the lower order correction terms of the mean curvature.

### 5.2. Curvature corrections for $H$ and its derivatives

By a slight variation of terms, one can also derive estimates for $H$ instead of $\frac{1}{2} H^{2}$. Alternatively one can proceed as follows. Recall that for functions $f, g \in C^{\infty}(\Sigma)$ with $g \neq 0$ we have the identity

$$
\nabla_{i, j}^{2} \frac{u}{v}=-\frac{1}{v^{2}}\left(\nabla_{i} v \nabla_{j} u+\nabla_{i} u \nabla_{j} v\right)+\frac{1}{v} \nabla_{i, j}^{2} u-\frac{u}{v^{2}} \nabla_{i, j}^{2} v+2 \frac{u}{v^{3}} u \nabla_{i} v \nabla_{j} v
$$

Using the a priori estimates for $H, \nabla H$ and $\nabla^{2} H$ as before, we find that

$$
\nabla_{i, j}^{2}\left(H^{-1} \operatorname{Ric}(\nu, \nu)\right)=H^{-1} \nabla_{i, j}^{2} \operatorname{Ric}(\nu, \nu)+O_{L^{2}}(1)
$$

so that equation 10 yields

$$
\nabla_{i, j}^{2}\left(H^{-1} \operatorname{Ric}(\nu, \nu)\right)=-\frac{1}{2} H \operatorname{Ric}(\nu, \nu) \gamma_{i j}+\frac{1}{2} H T_{i j}+O_{L^{2}}(1)
$$

Splitting into trace part and trace-free part we get

$$
\begin{equation*}
\Delta\left(H^{-1} \operatorname{Ric}(\nu, \nu)\right)=-\frac{3}{2} H \operatorname{Ric}(\nu, \nu)+\frac{1}{2} H \mathrm{Sc}+O_{L^{2}}(1) \tag{19}
\end{equation*}
$$

and

$$
\begin{equation*}
\left[\nabla^{2}\left(H^{-1} \operatorname{Ric}(\nu, \nu)\right)\right]_{i j}^{\circ}=\frac{1}{2} H \stackrel{\circ}{T}_{i j}+O_{L^{2}}(1) \tag{20}
\end{equation*}
$$

Let $v:=H-\frac{2}{3} H^{-1} \operatorname{Ric}(\nu, \nu)$. Combining equations (19) and (1), with the estimate from theorem 2.5 as in section 5.1 we find that

$$
\Delta v=O_{L^{2}}(1)
$$

Arguing as before, the Bochner identity implies:

$$
\begin{equation*}
\left\|\nabla^{2} v\right\|_{L^{2}} \leq C R \quad \text { and } \quad\|\nabla v\|_{L^{2}} \leq C R^{2} \tag{21}
\end{equation*}
$$

These considerations imply the following estimate.
Proposition 5.3. Assume that $(M, g)$ and $\Sigma$ are as in Proposition 5.1. Then

$$
\left\|\left(\nabla^{2} H\right)^{\circ}-\frac{1}{3} H \stackrel{\circ}{T}\right\|_{L^{2}} \leq C R(\Sigma) \quad \text { and } \quad\left\|\nabla H-\frac{2}{3} \omega\right\|_{L^{2}} \leq C R(\Sigma)^{2}
$$

Proof. The proof follows directly from the estimates (21) in combination with formulas (9) and (12).

### 5.3. Curvature corrections for $\AA$

To estimate the corrections of the curvature to $\AA$ recall the Simons-Identity on $\Sigma$ as in equation (5). In view of the a priori estimates from theorem 2.3 and the conventions in in section 4 , on surfaces as in theorem 2.3 this yields

$$
\begin{equation*}
\Delta \AA_{i j}=\left(\nabla^{2} H\right)_{i j}^{0}+\frac{1}{2} H^{2} \AA_{i j}+2 \nabla_{i} \omega_{j}-\operatorname{div} \omega \gamma_{i j}+O_{L^{\infty}}(R) \tag{22}
\end{equation*}
$$

In view of proposition 5.3 and equation (15) we thus infer

$$
\Delta \AA_{i j}=\frac{4}{3} H \stackrel{\circ}{T}_{i j}+\frac{1}{2} H^{2} \AA_{i j}+O_{L^{2}}(1)
$$

In view of the a priori estimates and equation (16) we find that

$$
\Delta\left(H^{-1} \stackrel{\circ}{T}\right)_{i j}=-\frac{1}{2} H \stackrel{\circ}{T}_{i j}+O_{L^{2}}(1)
$$

so that the tensor

$$
S_{i j}:=\stackrel{\circ}{A}_{i j}+\frac{4}{3} H^{-1} \stackrel{\circ}{T}_{i j}
$$

satisfies

$$
\begin{equation*}
\Delta S_{i j}=\frac{1}{2} H^{2} S_{i j}+O_{L^{2}}(1) \tag{23}
\end{equation*}
$$

Multiplying (23) by $S^{i j}$ and integrating by parts implies

$$
\int_{\Sigma}|\nabla S|^{2}+\frac{1}{2} H^{2}|S|^{2} \mathrm{~d} \mu \leq \int_{\Sigma}|S| \mathrm{d} \mu \leq \frac{1}{4} \int_{\Sigma} H^{2}|S|^{2} \mathrm{~d} \mu+\int_{\Sigma} H^{-2} \mathrm{~d} \mu
$$

Absorbing the first term on the right to the left yields the following estimate.
Proposition 5.4. Assume that $(M, g)$ and $\Sigma$ are as in Proposition 5.1. Then

$$
\left\|\AA+\frac{4}{3} H^{-1} \stackrel{\circ}{T}\right\|_{L^{2}} \leq C R(\Sigma)^{3} \quad \text { and } \quad\left\|\nabla \AA+\frac{4}{3} H^{-1} \nabla \stackrel{\circ}{T}\right\|_{L^{2}} \leq C R(\Sigma)^{2}
$$

## 6. Expansion of the metric

It is well known ${ }^{1}$ that the metric in normal coordinates has the expansion

$$
g_{\alpha \beta}(y)=\delta_{\alpha \beta}+\frac{1}{3} \operatorname{Rm}_{\alpha \mu \beta \nu} y^{\mu} y^{\nu}+O\left(|y|^{2}\right)
$$

Here we denote $\mathrm{Rm}_{\alpha \mu \beta \nu}=\operatorname{Rm}_{\alpha \mu \beta \nu}(0)$ and all other curvature quantities are evaluated at 0 as well. From this we calculate that

$$
g_{\alpha \beta, \mu \nu}(0)=\frac{1}{3}\left(\mathrm{Rm}_{\alpha \mu \beta \nu}+\mathrm{Rm}_{\alpha \nu \beta \mu}\right)
$$

The Christoffel symbols thus satisfy

$$
\begin{aligned}
& \Gamma_{\alpha \beta, \gamma}^{\nu}(0) \\
& \quad=\frac{1}{2} \delta^{\nu \mu}\left(g_{\alpha \mu, \beta \gamma}+g_{\beta \mu, \alpha \gamma}-g_{\alpha \beta, \mu \gamma}\right) \\
& \quad=\frac{1}{6} \delta^{\nu \mu}\left(\operatorname{Rm}_{\alpha \beta \mu \gamma}+\operatorname{Rm}_{\alpha \gamma \mu \beta}+\operatorname{Rm}_{\beta \alpha \mu \gamma}+\operatorname{Rm}_{\beta \gamma \mu \alpha}-\operatorname{Rm}_{\alpha \mu \beta \gamma}-\operatorname{Rm}_{\alpha \gamma \beta \mu}\right) \\
& \quad=\frac{1}{3} \delta^{\nu \mu}\left(\operatorname{Rm}_{\alpha \gamma \mu \beta}+\operatorname{Rm}_{\beta \gamma \mu \alpha}\right) \\
& \quad=-\frac{1}{3} \delta^{\nu \mu}\left(\operatorname{Rm}_{\alpha \gamma \beta \mu}+\operatorname{Rm}_{\beta \gamma \alpha \mu}\right)=-\frac{1}{3}\left(\operatorname{Rm}_{\alpha \gamma \beta}^{\nu}+\operatorname{Rm}_{\beta \gamma \alpha}^{\nu}\right) .
\end{aligned}
$$

From this we get that in normal coordinates

$$
\begin{aligned}
\Gamma_{\alpha \beta}^{\nu}(y) & =\Gamma_{\alpha \beta}^{\nu}(0)+\Gamma_{\alpha \beta, \gamma}^{\nu}(0) y^{\gamma}+O\left(|y|^{2}\right) \\
& =-\frac{1}{3}\left(\operatorname{Rm}_{\alpha \gamma \beta}^{\nu}+\operatorname{Rm}_{\beta \gamma \alpha}^{\nu}\right) y^{\gamma}+O\left(|y|^{2}\right) .
\end{aligned}
$$

[^0]This implies that for a constant vector field $b=b^{\nu} \partial_{\nu} \in \mathbf{R}^{3}$ we have

$$
\begin{equation*}
\nabla_{\alpha} b^{\nu}=\partial_{\alpha} b^{\nu}+\Gamma_{\alpha \beta}^{\nu} b^{\beta}=-\frac{1}{3}\left(\operatorname{Rm}_{\alpha \gamma \beta}^{\nu}(0)+\operatorname{Rm}_{\beta \gamma \alpha}^{\nu}(0)\right) y^{\gamma} b^{\beta}+O\left(|y|^{2}\right) \tag{24}
\end{equation*}
$$

and

$$
\begin{equation*}
\operatorname{div} b=\nabla_{\alpha} b^{\alpha}=-\frac{1}{3} \operatorname{Ric}_{\beta \gamma}(0) y^{\gamma} b^{\beta}+O\left(|y|^{2}\right) \tag{25}
\end{equation*}
$$

Terms like these will show up in the position estimates. Here we explicitly included the point at which to evaluate the curvature for later reference.

## 7. The position estimates revisited

The basic idea of the position estimates for small area constrained Willmore surfaces in [6] is to test the Euler-Lagrange-Equation

$$
\begin{equation*}
\delta_{f} \mathcal{W}(\Sigma)=\lambda \int_{\Sigma} f H \mathrm{~d} \mu \tag{26}
\end{equation*}
$$

with the function $f=H^{-1} g(b, \nu)$. The main result of [6] is the estimate

$$
\begin{equation*}
|\nabla \mathrm{Sc}(p)| \leq C R(\Sigma) \tag{27}
\end{equation*}
$$

where $p$ is a point with $\operatorname{dist}(p, \Sigma) \leq \operatorname{diam}(\Sigma)$ and $C$ is a constant depending only on $C_{B}$. With this choice of point, we have that $r(x)=\operatorname{dist}(p, x)$ for $x \in \Sigma$ is comparable to the area radius $R(\Sigma)$.

To improve this estimate further we have to carefully choose the center point $p$ of the above coordinates. The main result of the paper in this section is:

Theorem 7.1. Let $(M, g)$ be a 3-manifold with $C_{B}$-bounded geometry. Then there exist constants $a_{0} \in(0, \infty)$ and $C \in(0, \infty)$ depending only on $C_{B}$ with the following property. Let $\Sigma \subset M$ be a surface satisfying the EulerLagrange equation (1) for some $\lambda \in \mathbf{R}$, with $|\Sigma| \leq a_{0}$, and $\mathcal{W}(\Sigma) \leq 4 \pi+a_{0}$. Then there exists a point $p_{0} \in M$ such that
i) $\left|\operatorname{dist}\left(p_{0}, x\right)-R(\Sigma)\right| \leq C R(\Sigma)^{3}$ for all $x \in \Sigma$,
ii) with respect to normal coordinates $\psi: \mathcal{B}_{\rho}\left(p_{0}\right) \rightarrow B_{\rho}(0) \subset \mathbf{R}^{3}$ centered at $p_{0}$ we have

$$
\int_{\psi(\Sigma)} y^{\alpha} \mathrm{d} \mu_{g}(y)=0
$$

iii) and $\left|\nabla \mathrm{Sc}\left(p_{0}\right)\right| \leq C R(\Sigma)^{2}$.

Remark 7.2. Note that appealing to theorem 2.11, lemma 2.12 and estimate 2.3 , we automatically have that $\Sigma$ is $W^{2,2}$-close to the geodesic sphere of radius $R$ around $p_{0}$ in the following sense. Denote by $h_{R}: \mathbf{R}^{3} \rightarrow \mathbf{R}^{3}$ : $y \mapsto \frac{y}{R}$ the scaling vector field. Let $\Sigma_{R}:=h_{R}\left(\psi(\Sigma) \subset B_{\frac{\rho}{R}}(0)\right.$. Then there is a map $F: S^{2} \rightarrow \Sigma_{R}$, conformal with respect to metric on $\Sigma_{R}$ induced by the Euclidean metric such that

$$
\|F\|_{W^{2,2}\left(S^{2}\right)} \leq C R(\Sigma)^{2}
$$

For the proof of Theorem 7.1 assume that $\Sigma$ is as in the statement of Theorem 7.1 and that $a_{0}$ is chosen so small that all the estimates from sections 2 to section 6 are applicable. In particular, Lemma 3.1 gives a point $p_{0} \in M$ such that $\left|\operatorname{dist}\left(p_{0}, x\right)-R(\Sigma)\right| \leq C R(\Sigma)^{3}$ for all $x \in \Sigma$ and such that if $\psi: \mathcal{B}_{\rho}\left(p_{0}\right) \rightarrow B_{\rho}(0) \subset \mathbf{R}^{3}$ are normal coordinates at $p_{0}$ then $\int_{\psi(\Sigma)} y \mathrm{~d} \mu_{g}=0$. Then the first two assertions of the Theorem directly follow. The estimate for $\left|\nabla \operatorname{Sc}\left(p_{0}\right)\right|$ follows from the calculations in the remainder of this section. All these calculations are done in the normal coordinates centered at $p_{0}$.

Recall the splitting

$$
\begin{equation*}
\delta_{f} \mathcal{W}(\Sigma)=\delta_{f} \mathcal{U}(\Sigma)+\delta_{f} \mathcal{V}(\Sigma) \tag{28}
\end{equation*}
$$

that was used with (26) for the test function $f=H^{-1} g(b, \nu)$. Here $b \in \mathbf{R}^{3}$ is a constant vector in the normal coordinate neighborhood. The computations below use the same $f$ and the same splitting.

### 7.1. The right hand side

For $f=H^{-1} g(b, \nu)$ we have that

$$
\int_{\Sigma} f H \mathrm{~d} \mu=\int_{\Sigma} g(b, \nu) \mathrm{d} \mu=\int_{\Omega} \operatorname{div} b \mathrm{~d} V
$$

where $\Omega$ is the region enclosed by $\Sigma$. In the integral on the right, we replace the volume form of $g$ by the Euclidean volume form of the normal coordinates at $p_{0}$ and obtain an error of the form

$$
\left|\int_{\Omega} \operatorname{div} b \mathrm{~d} V-\int_{\Omega} \operatorname{div} b \mathrm{~d} V^{E}\right| \leq \operatorname{Vol}(\Omega) \sup _{\Omega}|\nabla b| R^{2} \leq C|\Sigma|^{3}
$$

since $|\nabla b|=O(R),\left|\mathrm{d} V-\mathrm{d} V^{E}\right|=O\left(R^{2}\right) \mathrm{d} V^{E}$ and $\operatorname{Vol}(\Omega)=O\left(R^{3}\right)$ by [6, Eq. (4.7)]. As usual we abbreviate $R=R(\Sigma)$. At this point we do not care about errors of the order $O\left(R^{5}\right)$ but want to compute the top order term which is $O\left(R^{4}\right)$. In section 6 we computed that

$$
\operatorname{div} b=-\frac{1}{3} \operatorname{Ric}_{\alpha \beta} y^{\alpha} b^{\beta}+O\left(R^{2}\right)
$$

Note that the volume integral of the error term is $O\left(R^{5}\right)$, and that Ric here is evaluated at $p_{0}$, the origin of the normal coordinates $y$. We thus get

$$
\begin{aligned}
\int_{\Omega} \operatorname{div} b \mathrm{~d} \mu & =-\frac{1}{3} \operatorname{Ric}_{\alpha \beta} b^{\beta} \int_{\Omega} y^{\alpha} \mathrm{d} \mu^{E}+O\left(R^{5}\right) \\
& =-\frac{1}{6} \operatorname{Ric}_{\alpha \beta} b^{\beta} \int_{\Sigma}|y|^{2}\left(\nu^{E}\right)^{\alpha} \mathrm{d} \mu^{E}+O\left(R^{5}\right)
\end{aligned}
$$

Here $\nu_{E}$ denotes the normal to $\Sigma$ with respect to the Euclidean metric in our coordinates. From Corollary 3.3 with $k=1, \rho_{1}=\rho_{2}=\frac{y^{\beta}}{R}$ and $\rho_{3}=\left(\nu^{E}\right)^{\alpha}$ it follows that

$$
\int_{\Sigma}\left(y^{\beta}\right)^{2}\left(\nu^{E}\right)^{\alpha} \mathrm{d} \mu^{E} \leq C R^{5}
$$

and after summation over $\beta$ we arrive at

$$
\left|\int_{\Omega} \operatorname{div} b \mathrm{~d} \mu\right| \leq C R^{5}
$$

or, in combination with the estimates above, with 26 and using Theorem 2.5 this gives

$$
\begin{equation*}
\left|\delta_{f} \mathcal{W}(\Sigma)\right| \leq C R^{5} \tag{29}
\end{equation*}
$$

Note that this improves the estimate from [6] by one power of $R$.

### 7.2. The variation of $\mathcal{U}(\Sigma)$

From [6] we have

$$
\begin{equation*}
\delta_{f} \mathcal{U}(\Sigma)=-\int_{\Sigma} 2\left\langle\AA, \nabla^{2} f\right\rangle+2 f\langle\AA \stackrel{\circ}{\Pi}\rangle+f H|\AA|^{2} \mathrm{~d} \mu \tag{30}
\end{equation*}
$$

and that for our choice $f=H^{-1} g(b, \nu)$ we have

$$
\begin{aligned}
\nabla_{i j}^{2} f= & -A_{i}^{k} A_{j k} f+H^{-1} g\left(\nabla_{i} b, e_{k}\right) A^{k j}-H^{-2} \nabla_{i} H g\left(b, e_{k}\right) A_{j}^{k} \\
& +\nabla_{i}\left(H^{-1} g\left(\nabla_{j} b, \nu\right)-H^{-2} \nabla_{j} H g(b, \nu)\right)
\end{aligned}
$$

Since

$$
A_{i}^{k} A_{j k}=\AA_{i}^{k} \AA_{j k}+H \AA_{i j}+\frac{1}{4} H^{2} g_{i j}
$$

and since the first and the last term give zero when contracted with $\AA$ we get that

$$
\begin{aligned}
\int_{\Sigma}\left\langle\nabla^{2} f, \AA\right\rangle \mathrm{d} \mu=\int_{\Sigma} & \operatorname{div} \AA^{j}\left(H^{-2} \nabla_{j} H g(b, \nu)-H^{-1} g\left(\nabla_{j} b, \nu\right)\right)+f H|\AA|^{2} \\
& -H^{-2} \nabla_{i} H g\left(b, e_{k}\right) A_{j}^{k} A^{i j}+H^{-1} g\left(\nabla_{i} b, e_{k}\right) A_{j}^{k} \AA^{i j} \mathrm{~d} \mu
\end{aligned}
$$

Plugging into we get that

$$
\begin{aligned}
\delta_{f} \mathcal{U}(\Sigma)=2 & \int_{\Sigma} \operatorname{div} \AA^{j}\left(H^{-1} g\left(\nabla_{j} b, \nu\right)-H^{-2} \nabla_{j} H g(b, \nu)\right)-\frac{3}{2} f H|\AA|^{2} \\
& +H^{-2} \nabla_{i} H g\left(b, e_{k}\right) A_{j}^{k} \AA^{i j}-H^{-1} g\left(\nabla_{i} b, e_{k}\right) A_{j}^{k} \AA^{i j}-f\langle\AA, \stackrel{\circ}{T}\rangle \mathrm{d} \mu .
\end{aligned}
$$

We shall only keep the top order parts of the first two terms in the second line. In view of the $L^{\infty}$ estimates for $\AA, H^{-1}$ and the $L^{2}$-estimates for $\nabla H$, we have that

$$
\begin{align*}
\delta_{f} \mathcal{U}(\Sigma)= & \int_{\Sigma} 2 \operatorname{div} \AA^{j}\left(H^{-1} g\left(\nabla_{j} b, \nu\right)-H^{-2} \nabla_{j} H g(b, \nu)\right)-3 f H|\AA|^{2}  \tag{31}\\
& \quad+H^{-1} \nabla_{i} H g\left(b, e_{j}\right) \AA^{i j}-g\left(\nabla_{i} b, e_{j}\right) \AA^{i j}-2 f\langle\AA, \stackrel{\circ}{T}\rangle \mathrm{d} \mu+O\left(R^{5}\right)
\end{align*}
$$

In view of the estimates in section 5 all the above terms can be replaced with their highest order parts. The error terms are then of order $O\left(R^{5}\right)$ or better. To be specific, we recall that to top order

$$
H^{-1} \approx \frac{R}{2}, \quad \AA \approx-\frac{4}{3} H^{-1} \stackrel{\circ}{T}, \quad \nabla H \approx \frac{2}{3} \omega, \quad \text { and } \quad \operatorname{div} \AA=\frac{4}{3} \omega .
$$

This yields that

$$
\begin{align*}
\delta_{f} \mathcal{U}(\Sigma)=\frac{2}{3} \int_{\Sigma} & 2 R \omega_{j} g\left(\nabla_{j} b, \nu\right)-\frac{2}{3} R^{2}|\omega|^{2} g(b, \nu)-R^{2} g(b, \nu)|\stackrel{\circ}{T}|^{2}  \tag{32}\\
& -\frac{1}{3} R^{2} \omega_{i} g\left(b, e_{j}\right) \grave{T}^{i j}+R g\left(\nabla_{i} b, e_{j}\right) \stackrel{\circ}{T}^{i j} \mathrm{~d} \mu+O\left(R^{5}\right)
\end{align*}
$$

Note that all the previous terms can be expanded into integrals that can individually be estimated using Corollary 3.3. Consider for example the first
term on the right of (32):

$$
\begin{aligned}
& \frac{4 R}{3} \int_{\Sigma} \sum_{j=1}^{2} \omega\left(e_{j}\right) g\left(\nabla_{e_{j}} b, \nu\right) \mathrm{d} \mu \\
& =\frac{4 R}{3} \int_{\Sigma} \sum_{\beta=1}^{3}\left(\operatorname{Ric}_{\alpha \beta} \nu^{\alpha} g_{\eta \mu} \nabla_{\beta} b^{\eta} \nu^{\mu}\right)-\operatorname{Ric}_{\alpha \beta} \nu^{\alpha} \nu^{\beta} g_{\eta \mu} \nabla_{\kappa} b^{\eta} \nu^{\mu} \nu^{\kappa} \mathrm{d} \mu
\end{aligned}
$$

After replacing $\nabla b$ using the expansion (24), along with

$$
g_{\eta \mu}=\delta_{\eta \mu}+O\left(R^{2}\right), \quad \text { and } \quad \operatorname{Ric}=\operatorname{Ric}\left(p_{0}\right)+O(R)
$$

and noting that the resulting error terms are of order $O\left(R^{5}\right)$ we can use Corollary 3.3 to see that the whole term is $O\left(R^{5}\right)$. Inspecting the other terms of (32) shows that they can be treated similarly. Indeed all the tangential contractions in these terms can be resolved as above and the remaining terms are products of an odd number of factors $\nu$ or $y / R$. To show the pattern note that

$$
\begin{aligned}
|\omega|^{2} & =\left|\operatorname{Ric}(\nu, \cdot)^{T}\right|^{2}=|\operatorname{Ric}(\nu, \cdot)|^{2}-\operatorname{Ric}(\nu, \nu)^{2} \\
& =\sum_{\beta=1}^{3} g_{\beta \kappa} \operatorname{Ric}_{\alpha \beta} \operatorname{Ric} \eta \kappa \nu^{\alpha} \nu^{\eta}-\left(\operatorname{Ric}_{\alpha \beta} \nu^{\alpha} \nu^{\beta}\right)^{2}
\end{aligned}
$$

Both terms on the right have an even number of factors $\nu$ so that multiplied with $g(b, \nu)$ in the second term on the right of (32) yields an odd number. The third term can be treated by computing with $\tau=\operatorname{tr} T=\operatorname{Sc}-\operatorname{Ric}(\nu, \nu)$ that

$$
\begin{aligned}
|\stackrel{\circ}{T}|^{2} & =|T|^{2}-\frac{1}{2} \tau^{2} \\
& =|\operatorname{Ric}|^{2}-2|\omega|^{2}-\operatorname{Ric}(\nu, \nu)^{2}-\frac{1}{2} \operatorname{Sc}^{2}-\operatorname{Sc} \operatorname{Ric}(\nu, \nu)-\frac{1}{2} \operatorname{Ric}(\nu, \nu)^{2}
\end{aligned}
$$

Note that all terms on the right contain an even number of factors $\nu$, so the third term in (32) is also done. The remaining two terms have a similar structure. We infer the estimate

$$
\begin{equation*}
\left|\delta_{f} \mathcal{U}(\Sigma)\right| \leq C R^{5} \tag{33}
\end{equation*}
$$

for the particular choice of $f$ above.

### 7.3. The variation of $\mathcal{V}(\Sigma)$

From [6, Section 4.3] we get that for the given choice of $f$ we have

$$
\begin{align*}
\delta_{f} \mathcal{V}(\Sigma)= & \int_{\Sigma}-G(b, \nu)-\frac{1}{2} g(b, \nu) \mathrm{Sc}+2 f\left\langle\AA, G^{T}\right\rangle \\
& -2 \omega\left(e_{i}\right)\left(H^{-1} g\left(\nabla_{e_{i}} b, \nu\right)+H^{-1} \AA_{i}^{j} g\left(b, e_{j}\right)\right.  \tag{34}\\
& \left.-H^{-2} \nabla H g(b, \nu)\right) \mathrm{d} \mu .
\end{align*}
$$

As in section 7.2 we can estimate

$$
\begin{aligned}
& \left|\int_{\Sigma} 2 f\left\langle\AA, G^{T}\right\rangle-2 \omega\left(e_{i}\right) H^{-2}\left(H g\left(\nabla_{e_{i}} b, \nu\right)+H \AA_{i}^{j} g\left(b, e_{j}\right)-\nabla H g(b, \nu)\right) \mathrm{d} \mu\right| \\
& \quad \leq C R^{5} .
\end{aligned}
$$

To see this, use $\nabla H=\frac{2}{3} \omega+O_{L^{2}}(R)$ and $\AA=-\frac{4}{3} H^{-1} \stackrel{\circ}{T}+O_{L^{2}}\left(R^{2}\right)$ from Propositions 5.3 and 5.4. The estimate then follows by inspection as in section 7.2,

Furthermore, as in [6, Section 4.3] let $X$ be the vector field on $\mathcal{B}_{\rho}\left(p_{0}\right) M$ such that $g(X, Y)=G(b, Y)$ for all vector fields $Y$ on $\mathcal{B}_{\rho}\left(p_{0}\right)$. Then $\operatorname{div}_{M} X=\langle G, \nabla b\rangle$ since $G$ is divergence free. Let $\Omega \subset \mathcal{B}_{\rho}\left(p_{0}\right)$ enclosed by $\Sigma$ and recall from [6, Section 4.3] that $\operatorname{Vol}(\Omega) \leq C R^{3}$. Compute

$$
\begin{aligned}
& \int_{\Sigma} G(b, \nu) \mathrm{d} \mu=\int_{\Omega} \operatorname{div}_{M} X \mathrm{~d} V=\int_{\Omega}\langle G, \nabla b\rangle \mathrm{d} V=\int_{\Omega} G_{\alpha \beta} \nabla_{\kappa} b^{\alpha} g^{\beta \kappa} \mathrm{d} V \\
& =-\frac{1}{3} G_{\alpha \beta}\left(p_{0}\right) \delta^{\beta \kappa}\left(\operatorname{Rm}_{\kappa \eta \mu}^{\alpha}\left(p_{0}\right)+\operatorname{Rm}_{\mu \eta \kappa}^{\alpha}\left(p_{0}\right)\right) b^{\mu} \int_{\Omega} y^{\kappa} \mathrm{d} V^{E}+O\left(R^{5}\right)
\end{aligned}
$$

Here we used equation (24) in the last step and replaced all curvature quantities by their values at $p_{0}$. Also the integration is now with respect to the Euclidean volume form $\mathrm{d} V^{E}$. The value of the constant in front of the integral is not important for the following. For $\kappa \in\{1,2,3\}$ consider the vector field $Y:=\frac{1}{4} y^{\kappa} y$. Then $\operatorname{div}_{E} Y=y^{\kappa}$ and thus using Stokes in the first equality and Corollary 3.3 in the estimate.

$$
\begin{equation*}
\int_{\Omega} y^{\kappa} \mathrm{d} V^{E}=\int_{\Sigma} y^{\kappa}\left\langle y, \nu^{E}\right\rangle_{E} \mathrm{~d} \mu^{E}=O\left(R^{5}\right) \tag{35}
\end{equation*}
$$

To treat the remaining term, we consider the vector field $X=\operatorname{Sc} b$ as in 6, Section 4.3]. Then

$$
\int_{\Sigma} g(b, \nu) \operatorname{Sc} \mathrm{d} \mu=\int_{\Omega} \operatorname{div}_{M} X \mathrm{~d} V=\int_{\Omega} g(b, \nabla \mathrm{Sc})+\mathrm{Sc}_{\operatorname{div}}^{M} \text { } b \mathrm{~d} V
$$

with $\Omega$ as above. Using $g=g^{E}+O\left(R^{2}\right), \mathrm{Sc}=\operatorname{Sc}\left(p_{0}\right)+O(R)$,

$$
\nabla_{\beta} \mathrm{Sc}=\nabla_{\beta} \mathrm{Sc}(0)+\nabla_{\beta, \kappa}^{2} \mathrm{Sc}(0) y^{\kappa}+O\left(R^{2}\right)
$$

and equation (25) for the expansion of $\operatorname{div} b$, we get

$$
\begin{aligned}
\int_{\Sigma} g(b, \nu) \operatorname{Sc~d} \mu= & \int_{\Omega} g^{E}\left(b, \nabla \operatorname{Sc}\left(p_{0}\right)\right) \mathrm{d} V^{E} \\
& +\left(b^{\alpha} \nabla_{\alpha \kappa}^{2} \operatorname{Sc}-\frac{1}{3} \operatorname{Sc}\left(p_{0}\right) \operatorname{Ric}_{\beta \kappa}\left(p_{0}\right) b^{\beta}\right) \int_{\Omega} y^{\kappa} \mathrm{d} V^{E}+O\left(R^{5}\right)
\end{aligned}
$$

In view of (35) this gives

$$
\int_{\Sigma} g(b, \nu) \operatorname{Sc~} \mathrm{d} \mu=\operatorname{Vol}(\Omega) g\left(b, \nabla \operatorname{Sc}\left(p_{0}\right)\right)+O\left(R^{5}\right)
$$

In combination with the above, we arrive at the estimate

$$
\begin{equation*}
\left|\delta_{f} \mathcal{V}(\Sigma)+\frac{1}{2} \operatorname{Vol}(\Omega) g\left(b, \nabla \operatorname{Sc}\left(p_{0}\right)\right)\right| \leq C R(\Sigma)^{5} . \tag{36}
\end{equation*}
$$

### 7.4. The conclusion

From the splitting (28), estimates (29), (33), and (36) we arrive at

$$
\operatorname{Vol}(\Omega)\left|g\left(b, \nabla \operatorname{Sc}\left(p_{0}\right)\right)\right| \leq C R(\Sigma)^{5} .
$$

Since $b \in \mathbf{R}^{3}$ is arbitrary and since $\operatorname{Vol}(\Omega) \geq C^{-1} R(\Sigma)^{3}$ by [6, Eq. (4.7)] this gives the claimed estimate:

$$
\left|\nabla \mathrm{Sc}\left(p_{0}\right)\right| \leq C R(\Sigma)^{2}
$$

This concludes the proof of Theorem 7.1.

## 8. The proof of Corollary 1.5

Corollary 8.1. Let $(M, g)$ be a compact three dimensional Riemannian manifold with $C_{B}$ bounded geometry. Let

$$
Z:=\{x \in M \mid \nabla \mathrm{Sc}(x)=0\}
$$

and assume that the Hessian $\operatorname{Hess} \operatorname{Sc}(x)$ is non-degenerate for every $x \in Z$.

Then there exists an $a_{0}$ depending only on $(M, g)$ such that for every surface $\Sigma$ that satisfies the Euler-Lagrange equation (1) for some $\lambda$, with $|\Sigma| \leq a_{0}$ and $\mathcal{W}(\Sigma) \leq 4 \pi+a_{0}$ the region enclosed by $\Sigma$ intersects $Z$ in a single point.

Proof. Since $M$ is compact and all critical points of Sc are non-degenerate, the set $Z$ is discrete. Let

$$
\rho_{0}:=\frac{1}{2} \min \{\operatorname{dist}(x, y) \mid x \neq y \in Z\}
$$

For $\rho \in\left(0, \rho_{0}\right)$ let

$$
Z_{\rho}:=\{x \in M \mid \operatorname{dist}(x, Z)<\rho\} .
$$

For $r \in(0, \infty)$ let

$$
G_{r}:=\{x \in M| | \nabla \operatorname{Sc}(x) \mid \leq r\}
$$

By the compactness of $M$ and since Sc is a Morse function, there exist $r_{0} \in(0, \infty)$ and $c \in(0, \infty)$ such that for every $r \in\left(0, r_{0}\right)$ we have $G_{r} \subset Z_{c r}$.

Let $a_{0}$ and $C$ be the constants from Theorem 7.1 applied to $(M, g)$. By decreasing $a_{0}$, we can assume that in addition to the assertion of Theorem 7.1, we also have that $C R(\Sigma)^{2} \leq r_{0}$ and that $\operatorname{diam}(\Sigma)<\rho_{0}$ whenever $\Sigma$ satisfies the assumption of this Lemma with the chosen $a_{0}$.

Let $\Sigma$ be such a surface and let $R=R(\Sigma)$. Denote by $\Omega$ the open region enclosed by $\Sigma$. For $s \in(0, \infty)$ denote

$$
\Omega_{s}:=\{x \in \Omega \mid \operatorname{dist}(x, \Sigma)>s\} .
$$

Then $\Omega_{s}$ is an open subset of $\Omega$.
Let $p_{0}$ be the point from Theorem 7.1. Then $\left|\operatorname{dist}\left(p_{0}, x\right)-R\right| \leq C R^{3}$ for all $x \in \Sigma$ and $\left|\nabla \mathrm{Sc}\left(p_{0}\right)\right| \leq C R^{2}$. The first estimate shows that $p_{0} \in \Omega$, in fact $p_{0} \in \Omega_{\frac{3}{4} R}$ if we choose $a_{0}$ sufficiently small.

Let $\sigma: \stackrel{4}{=}\left|\nabla \mathrm{Sc}\left(p_{0}\right)\right|$. Then $\sigma \leq C R^{2}$ so that $p_{0} \in G_{\sigma} \subset Z_{c \sigma}$. This implies that there exists a point $p_{1} \in Z$ such that $p_{1} \in \Omega_{\frac{3}{4} R-c \sigma} \subset \Omega_{\frac{3}{4} R-C R^{2}}$. By choosing $a_{0}$ and thus $R$ smaller again, we can ensure that $\frac{3}{4} R-\stackrel{4}{C} R^{2} \geq \frac{R}{2}$, so that $Z \cap \Omega_{\frac{R}{2}} \neq \emptyset$. Since $\operatorname{diam}(\Sigma)<\rho_{0}$ we know that $\Omega_{\frac{R}{2}} \subset \Omega \subset B_{\rho_{0}}\left(p_{0}\right)$ and by the choice of $\rho_{0}$ the ball $B_{\rho_{0}}\left(p_{0}\right)$ can intersect $Z$ in at most one point.

From the expansion of the Willmore energy in [7. Corollary 5.6] we know that the minimizers $\Sigma_{a}^{\min }$ from Theorem 1.1 concentrate near the maxima of the scalar curvature of $M$. Thus a slight variant of the proof of Corollary 8.1 yields the proof of Corollary 1.6.

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[^0]:    ${ }^{1}$ We use the convention for the curvature tensor from [8, Section 2] that is $\operatorname{Rm}_{\alpha \beta \gamma \nu}=\left\langle\left(\nabla_{\alpha} \nabla_{\beta}-\nabla_{\beta} \nabla_{\alpha}\right) \partial_{\gamma}, \partial_{\nu}\right\rangle$ and $\operatorname{Rm}_{\alpha \beta \gamma}^{\nu}=g^{\nu \mu} \operatorname{Rm}_{\alpha \beta \gamma \mu}$.

